Fast Deblurring Method for Computed Tomography Medical Images Using a Novel Kernels Set

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Abstract

Medical images such as computed tomography (CT) are degraded by different types of blur due to the imperfect resolution of the imaging system, data loss at the acquisition time and other technical reasons. The fastest way to deblur an image is by convolving a special kernel to the corrupted image. Laplacian kernels are famous and widely used in this field, but the issue is only few kernels are presented. This paper is trying to simulate the blur problem using various types of blur and attempt to restore the degraded images by using twenty novel kernels. Moreover, these kernels were tested with five types of blur that are: Average, Box, Gaussian, Pillbox and Atmospheric turbulence blur to determine which type of blur is suitable to be employed with kernels the most. The accuracy of the experimental results is measured with five diverse methods along with the success and the failure ratios. Finally, these kernels are applied to naturally degraded images obtained from different CT imaging systems.

Keywords: Computed Tomography (CT), Deblurring CT medical images, Sharpening kernels, Image deblurring, Average blur, Box blur, Gaussian blur, Pillbox blur, Atmospheric turbulence blur, Accuracy measurement methods

1. Introduction

Digital images are matrixes of numbers employed to show vital information. The methods of capturing and recording these images are diverse; therefore, the probabilities of having errors or degradations throughout the procedure of capturing and recording images are also increased [1]. The degradations that affect an image are noise, contrast imperfections, and blur [2]. This paper handles the blur degradation only. Blurring is uneasy to evade in every imaging device [18]. Blurry images are formed by convolving the original images with the point-spread function (PSF) [19]. CT medical images are known to be affected by blur [22] [7] [9]. Blur degrades CT images by the reasons of Gaussian noise [3], employing a denoising procedure on the degraded image [4], imperfect resolution of the imaging system [7], losing information throughout the acquisition process [8], and employing low-pass filters for reducing noise leads to blur amplification [9]. Blur has diverse types such as, atmospheric turbulence [12], Average [15], Box [17], Gaussian [14], Pillbox blur [16], and so on. Image deblurring is an essential topic in the area of image processing. The deblurring process results in sharpened details, better image quality and visualization [19]. The image restoration is a vital phase to recover images from their degradations; these techniques are considered as direct techniques when the outcome is formed in a one-step mode. Consistently, it's considered as indirect techniques when the outcome is acquired with a number of iterations. Famous restoration methods such as Wiener Filtering and Richardson-Lucy are examples of direct and indirect methods. The problems with these techniques are the essential need for the point-spread function (PSF), and determining the sufficient iterations required to restore the image [15]. Therefore, the use of kernels is more suitable, because determining the PSF and/or the number of iteration is not required.

2. Deblurring Procedure

Using kernels to deblur images is very simple. The basic concept is to convolve the kernel with the blurry image to obtain a sharper image. It takes one mathematical operation only, and it's fast and reliable. Suppose the degraded image is (D), the kernel is (K), and the convolution process is (\otimes), the restored image (R) can be described as the subsequent:

$\mathbf{R} = \mathbf{D} \otimes \mathbf{K}$

Laplacian kernels are well-known in the sharpening field. The problem is it contains only few sets of kernels. Therefore, the process of sharpening cannot be tuned well; the kernels either sharpen more or less than the desired amount. Thus, more sets of Kernels are demanded. This paper presents twenty novel kernels to tune and get the exact sharpening amount. The new kernels are:



Each of the above kernels has a different sharpening amount depending on the type of the blur and the blur volume. Using the correct kernel would grant the image a better and precise sharpening amount. Therefore, all the kernels would be tested with the five types of blur that are mentioned earlier, and the accuracy of the resulted image would be measured in different measurement techniques.

3. Accuracy Measurement Methods

Calculating the precision of the resulted image to the original one is considered as an important step. Therefore, traditional measuring techniques are utilized such as, Peak Signal to Noise Ratio (PSNR), Root Mean Square Error (RMSE), and Signal to Noise Ratio (SNR). PSNR computes the peak error. Reasonably, a greater rate of PSNR is better since it shows that the ratio of Signal to Noise is higher. In this method, the 'signal' is the reference image, and the 'noise' is the error in restoration [11]. Greater values for SNR and PSNR refer to a minor alteration between the original image and the restored image. The key benefit of these methods is its calculation simplicity. The Root Mean Square Error (RMSE) is the square root of the mean square error (MSE), lesser values of RMSE refers to a lower deference to the referenced image, and that leads to a better-quality image [5]. The equations of PSNR [6], SNR [5] and RMSE [5] are:

$$PSNR = 20 \log_{10} \frac{255}{\sqrt{MSE}}$$
$$SNR = 10 \log_{10} \frac{\sum_{i=1}^{n} \sum_{j=1}^{m} (A_{ij})^{2}}{\sum_{i=1}^{n} \sum_{j=1}^{m} (A_{ij} - B_{ij})^{2}}$$
$$RMSE = \sqrt{MSE} = \sqrt{\frac{1}{MN} \sum_{i=1}^{n} \sum_{j=1}^{m} (A(i,j) - B(i,j))^{2}}$$

Where: (A) is the referenced image; (B) is the restored image; (MN) is the height and width of the image. Furthermore, another two methods were used to measure the accuracy of the resulted image, such as Improvement in Signal-to-Noise Ratio (ISNR) and Universal Image Quality Index (UIQI). The ISNR is usually employed to measure the quality of the image in any imaging device. It's a powerful tool because it involves the reference image, the corrupted image and the restored image in the quality measurement process [20]. The equation of ISNR is [20]:

$$ISNR = 10 \log \frac{MSE(S, C)}{MSE(S, \hat{S})}$$

Where, (S) is the reference image, (C) is the corrupted image, (\hat{S}) is the restored image. Similarly, the UIQI measurement of the quality between the reference and result images is separated into three diverse assessments: luminance, contrast, and structural comparisons [21]. The equation of UIQI is [21]:

$$\text{UIQI} = \frac{4 \,\mu_{x} \mu_{y} \,\sigma_{xy}}{\left(\mu_{x}^{2} + \mu_{y}^{2}\right) \left(\sigma_{x}^{2} + \sigma_{y}^{2}\right)}$$

Where, (σ_x^2) is the variance of (μ_x) , (σ_y^2) is the variance of (μ_y) , (σ_{xy}) is the covariance of (μ_x, μ_y) , $(\mu_x) = \{x_1 \dots x_n\}$ and $(\mu_y) = \{y_1 \dots y_n\}$ [13].

4. Experimental Results

To determine which type of blur kernels can restore, an experiment has been conducted on five types of blur, namely are: Average, Box, Gaussian, Pillbox and Atmospheric turbulence blur. The degraded images of different types of blur are demonstrated in Figure 1. Consequently, the twenty kernels were applied to each type of blur. The accuracy was measured to recognize the restoration capability for every kernel and similarly to know which type of blur kernels can deblur the best. This paper will show only three results for each type of blur, and they are the worst, average and the finest result. The results of the experiment are illustrated in Figures 2, 3, 4, 5 and 6. The accuracy measurements along with the success and failure ratios are clarified in the subsequent tables.



Figure 1. CT images from left to right, top to bottom: original image, image degraded by average blur [5x5] with PSNR (22.9296), image degraded by Box Blur R=2 with PSNR (22.9235), image degraded by Gaussian Blur R=2 with PSNR (21.7893), image degraded by Pillbox Blur R=2 with PSNR (25.6318), and image degraded by Atmospheric turbulence Blur k=0.002 with PSNR (23.8106)



Figure 2. (Average Blur Restoration) Images from left to right: The worst result by K18, the average result by K15, and the best result by K11



Figure 3. (Box Blur Restoration) Images from left to right: The worst result by K18, the average result by K20, and the best result by K11



Figure 4. (Gaussian Blur Restoration) Images from left to right: The worst result by K1, the average result by K7, and the best result by K20



Figure 5. (Pillbox Blur Restoration) Images from left to right: The worst result by K18, the average result by K6, and the best result by K11



Figure 6. (Atmospheric Turbulence Blur Restoration) Images from left to right: The worst result by K18, the average result by K14, and the best result by K8

Table 1	K1	K2		K.	3	K4			K5	K6		K7	K8	K9	K10
PSNR	23.2393	22.35	20	22.62	203	23.	23.2125).7704	21.6838		22.7070	22.6925	23.0789	18.3453
SNR	12.2076	11.32)3	11.58	11.5886		12.1809		.7388	10.6522		11.6754	11.6608	12.0473	7.3137
RMSE	0.0689	0.076	3	0.07	0.0740		0.0691		.0915	0.0824		0.0732	0.0733	0.0702	0.1210
ISNR	-0.3096	0.577	7	0.30	94	-0.1	2829 2		.1592	1.2458		0.2226	0.2372	-0.1493	4.5843
UIQI	0.6734	0.650	6	0.65	92 0		.6658 (.6221	0.6315		0.6502	0.6611	0.6688	0.5670
K11	K12	K13	K	14	K	K15		6	K17	K18		K19	K20		
23.3943	19.1059	17.7151	20.8	8183	19.4	744	23.02	85	23.2756	5 16.9594	4	22.3912	19.7919		
12.3626	8.0743	6.6835	9.7	866	8.44	428	11.99	68	12.2440	5.9278		11.3596	8.7603	Average B	lur Filter
0.0677	0.1108	0.1301	0.0	910	0.10	0.070)6	0.0686	0.1419		0.0759	0.1024	Size [5x5]
-0.4646	3.8237	5.2145	2.1	114	3.45	552	-0.098	88	-0.3460	5.9702	2	0.5384	3.1377		
0.6678	0.5906	0.5492	0.6	i099	0.58	844	0.658	31	0.6706	0.5403		0.6518	0.5944		

Table 1. The Accuracy	/ Measurement with	the Average Blue
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Table 2. The Accuracy Measurement wi	εη τηθ	вох	Blur
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Table 2	K1]	K 2	K	3]	K4		K5		K6	K7	K8	K9	K10
PSNR	23.2952	22.	6155	22.7	237	23.	3338 2		21.1367		2.1038	23.0426	22.8377	23.3015	18.8177
SNR	12.2636	11.	5838	11.6	920	12.	12.3022		10.1051		1.0722	12.0110	11.8061	12.2699	7.7861
RMSE	0.0684	0.0	0740	0.07	731	0.0	0.0681		0.0877		0.0785	0.0704	0.0721	0.0684	0.1146
ISNR	-0.3717	0.1	080	0.19	998	-0.	4103	1	.7868	0	0.8197	-0.1191	0.0858	-0.3780	4.1058
UIQI	0.7043	0.0	872	0.69	931	31 0.7		0	.6626	0	0.6723	0.6898	0.6962	0.7044	0.6127
K11	K12	K13		K14	K	15 K1		6	K17		K18	K19	K20		
23.5467	19.5386	18.263	3 2	1.3926	20.0)720	23.38	09	23.317	5	17.4440	22.7098	20.2472		
12.5151	8.5070	7.2317	10	0.3610	9.0	403	12.34	.92	12.2858	8	6.4123	11.6782	9.2156	Box Blı	rR=2
0.0665	0.1055	0.1221	0	.0852	0.0	992	0.067	78	0.0683	3	0.1342	0.0732	0.0972		
-0.6232	3.3849	4.6602	1	.5309	2.8	516	16 -0.457		-0.3939	9	5.4796	0.2137	2.6763		
0.7060	0.6353	0.5973	C	.6551	0.6	322	0.700)4	0.7019)	0.5884	0.6933	0.6377		

Table 3	K1		K2		K	3	ŀ	X4		K5		K6	K7	K8	K9	K10
PSNR	22.2999		25.326	8	24.02	202	22.	22.6349		26.1804		5.7020	24.4101	24.7403	22.7286	26.2791
SNR	11.2683		14.295	1	12.9	885	11.	11.6032		15.1487		4.6704	13.3784	13.7086	11.6969	15.2475
RMSE	0.0767		0.0542	2	0.06	29	0.0)738	0	.0491	(0.0519	0.0602	0.0579	0.0730	0.0485
ISNR	-0.5107		-3.537	5	-2.23	309	-0.3	8456	-4	.3911	-,	3.9127	-2.6208	-2.9510	-0.9393	-4.4898
UIQI	0.6660		0.7274	1	0.69	16	0.6	5756	0	.7361	().7344	0.7157	0.7120	0.6845	0.7283
K11	K12	K	K13	K	K14	K	K15 K1		6	K17		K18	K19	K20		
23.6244	26.2399	25.	.8424	25.	.8024	26.4	414	24.12	39	23.0159)	25.3188	24.8368	26.4529	<i>.</i> .	D/
12.5928	15.2083	14.	.8108	14.	7707	15.4	.098	13.09	23	11.9842	2	14.2872	13.8051	15.4212	Gaussia	in Blur
0.0659	0.0488	0.0	0510	0.0	0513	0.04	476	0.062	22	0.0707		0.0542	0.0573	0.0476	R=	:2
-1.8351	-4.4507	-4.0	.0531	-4.	0131	-4.6	521	-2.334		-1.2266	5	-3.5295	-3.0475	-4.6636		
0.6922	0.7289	0.7	7202	0.7	7321	0.73	334	0.708	30	0.6735		0.7142	0.7158	0.7368		

 Table 3. The Accuracy Measurement with the Gaussian Blur

Table 4. The Accuracy Measurement with the Pillbox Blur

Table 4	K1		K2		K.	3	K4			K5	K	6	K7	K8	К9	K10
PSNR	26.4307		23.6539)	25.00)98	25.	25.7808 2		0.6342	21.9	598	24.0466	24.8942	25.5266	16.8999
SNR	15.3991		12.6223	3	13.97	13.9782		14.7492		9.6026		282	13.0149	13.8626	14.4950	5.8682
RMSE	0.0477		0.0657		0.0562		0.0514		0.	.0930	0.07	98	0.0628	0.0569	0.0529	0.1429
ISNR	-0.7990		1.9779		0.62	19	-0.	.1491 4		.9976	3.67	20	1.5852	0.7376	0.1052	8.7319
UIQI	0.7562		0.6932		0.72	15	0.7	0.7345		.6494	0.66	531	0.6989	0.7179	0.7358	0.5749
K11	K12	K	13	K	514	K	15	K1	K16 K		J	K18	K19	K20		
27.2989	18.1527	16.0)221	20.3	2636	18.3	563	25.4993		27.2304	15	.2404	24.2244	18.7791		
16.2673	7.1210	4.99	905	9.2	2320	7.32	246	14.46	76	16.1988	3 4.	.2088	13.1928	7.7474	Pillbo:	x Blur
0.0432	0.1237	0.15	581	0.0	0970	0.12	208 0.053		31	0.0435	0.	.1730	0.0615	0.1151	R=	-2
-1.6671	7.4791	9.60	097	5.3	3682	7.27	755	0.132	25	-1.5986	10	.3914	1.4074	6.8527		
0.7460	0.6080	0.55	527	0.6	5327	0.59	970	0.716	57	0.7561	0.	.5434	0.7033	0.6091		

Table 5. The Accuracy Measurement with the Atmospheric Turbulence Blur

Table 5	K1	K2	K	3	ŀ	K4		K5	K6	K7	K8	K9	K10
PSNR	24.7109	27.010	67 26.3	3800	24.	9053	24	.5671	25.6063	26.1963	27.3565	24.9290	20.6316
SNR	13.6793	15.98	50 15.3	8483	13.	3.8736		.5355	14.5746	15.1647	16.3248	13.8973	9.6000
RMSE	0.0581	0.044	6 0.0	480	0.0	.0569 (0591	0.0524	0.0490	0.0429	0.0567	0.0930
ISNR	-0.9003	-3.206	50 -2.5	693	-1.0)946	-0.	.7565	-1.7956	-2.3857	-3.5458	-1.1183	3.1790
UIQI	0.6926	0.702	8 0.6	980	0.6	5909 0		6694	0.6797	0.6927	0.7106	0.6943	0.6054
K11	K12	K13	K14	K	15 K1		16 K17		K18	K19	K20		
26.9675	21.9256	19.6058	23.8448	22.1	1068	26.76	536 25.974		18.7646	26.8626	22.6590	Atmos	spheric
15.9359	10.8939	8.5741	12.8132	11.0)752	15.73	320	14.9429	7.7330	15.8309	11.6273	Turbule	nce Blur
0.0448	0.0801	0.1046	0.0642	0.0	785	785 0.045		0.0503	0.1153	0.0454	0.0736	k-(002
-3.1568	1.8851	4.2049	-0.0342	1.7	038	38 -2.9530		-2.1639	5.0460	-3.0519	1.1517	κ-υ	.002
0.7076	0.6329	0.5843	0.6535	0.6	247	0.70	18	0.6988	0.5749	0.7005	0.6366		

After measuring the accuracy, the success and failure ratios should be determined depending on the results illustrated in the above tables. The ratio will be computed depending on PSNR results, and it must be calculated for each blur category to identify the type of blur that can be restored efficiently using kernels. The ratios can be determined as the subsequent:

Success ratio = $\frac{SK}{TNK} * 100\%$ Failure ratio = $\frac{FK}{TNK} * 100\%$

Where, (SK) represents the number of kernels that scored a higher PSNR value than the PSNR value of the corresponding blurry image in Figure 1. (FK) represents the number of kernels that scored lower or equal PSNR values. (TNK) represents the total number of kernels. The success and failure ratios are illustrated in the subsequent table.

Blur Type	SK	FK	TNK	Success ratio %	Failure ratio %
Average	6	14	20	30%	70%
Box	7	13	20	35%	65%
Gaussian	20	0	20	100%	0%
Pillbox	4	16	20	20%	80%
Atmospheric Turbulence	14	6	20	70%	30%
Overall Success Ratio	51	l%	Overall	Failure Ratio	49%

Table 6. The Success and Failure Ratios According to PSNR Statistics

5. Discussion

This paper proves lots of new concepts that they are: all the mentioned five types of blur can be restored using kernels but with diverse ratios depending on the type of blur and the blur density. Table 6 proves that the total success ratio is 51% of kernels to sharpen five types of blur. Furthermore, the behavior of the average and box blur is nearly the same. This inference has been inspired by comparing the results of the accuracy measurement techniques and the success and failure ratios between the two types of blur. Besides, the lowest ratio of success can be seen in the Pillbox blur, but still it has a reasonably high PSNR value among the restored images in the five types of blur, and that leads to a fact that the Pillbox blur can be restored efficiently but with certain types of kernels only. However, the atmospheric turbulence blur shows promising results to be deblurred with kernels when it gave a 70% succession ratio, and it gave the uppermost PSNR with the lowest RMSE values and that point to a fact that this type of blur can be restored efficiently but with certain be restored efficiently but with precise type of kernels only. Lastly, the Gaussian blur is the most suitable type of blur to be restored with kernels due to its succession ratio that gave a 100% with 0% failure.

6. Applying Kernels to Naturally Degraded CT Images

In this section, several naturally degraded CT images were selected to be restored using the proposed kernels. Figures 7, 8, 9, and 10 illustrate the CT images and their restored versions.



Figure 7. Images From left to right: original image, restored by K11, restored by K2, restored by K13

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Figure 8. Images from left to right: original image, restored by K20



Figure 9. Images from left to right: original image, restored by K14



Figure 10. Images from left to right: original image, restored by K18

7. Conclusion

In ending words, kernels are the fastest way to restore blurry images, because only one mathematical operation is involved and no prior knowledge about the PSF is required. Furthermore, kernels can successfully restore the mentioned five types of blur. Likewise, the finest restored images by kernels are images blurred with atmospheric turbulence and Pillbox blur, although the fact that these types of blur have relatively large failure ratios, especially the Pill box blur. The best type of blur that can be used with kernels is the Gaussian blur because of its 100% success ratio. Moreover, the average blur and the Box blur have a reasonably similar behavior due to the converged results between them. As a final point, if the CT medical images are degraded with one of five types of blur mentioned earlier, they can be restored by utilizing the novel kernels set presented in this paper.

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